

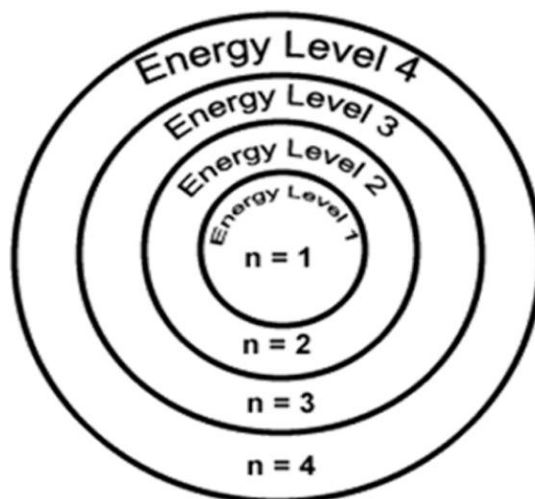
## ATOMIC MODEL

- See Atomic Model Timeline worksheet for specifics.

## ELECTRONS

- **Quantum Mechanical (QM) Model**- This is the currently accepted model of the atom.
  - Erwin Schrödinger wrote an equation which describes \_\_\_\_\_  
\_\_\_\_\_.
  - These locations are not definite because of the Heisenberg Uncertainty Principle.
- Each of the following terms gives a *more specific* description of where an electron *probably* is.
 

<b>In Chemistry...</b>	<b>In CB South, for example...</b>
○ Energy level, $n$	
○ Sublevel, $l$	
○ Orbital, $m_l$	
○ Spin, $s$	
- **Energy Levels,  $n$** -
  - An electron may NOT be found \_\_\_\_\_  
\_\_\_\_\_.
  - Higher  $n$  = higher energy (typically)
  - To determine how many electrons fit into a given energy level, use this **formula**: \_\_\_\_\_
  - The maximum number of electrons is **32**.
  - Electrons will occupy the \_\_\_\_\_ first.
- **Sublevels (subshells),  $l$** -
  - Energy levels contain \_\_\_\_\_ consisting of \_\_\_\_\_ (shapes) where there is a \_\_\_\_\_ probability of finding an \_\_\_\_\_.
  - Orbitals can hold up to \_\_\_\_\_ electrons.
  - Sublevels hold the orbitals and can hold 1, 3, 5, or 7 orbitals.



	<b><math>s</math> Sublevel</b>	<b><math>p</math> Sublevel</b>	<b><math>d</math> Sublevel</b>	<b><math>f</math> Sublevel</b>
<b>Shape</b>				
<b>Appears</b>				
<b># of Orbitals</b>				
<b>Capacity</b>				

## ELECTRON CONFIGURATIONS

- Writing Electron Configuration NOTES

Electron Configuration	Shorthand Configuration	Orbital Notation
H		
He		
Li		
Be		
B		
C		
Cl		
Ti		
Kr		

Electron Configuration	Shorthand Configuration	Orbital Notation
<b>O</b>		
<b>Ne</b>		
<b>Al</b>		
<b>Na</b>		
<b>Br</b>		
<b>Zr</b>		
<b>Ba</b>		
<b>Rn</b>		

## SHORTHAND NOTATION

- Steps to write in *shorthand* electron configuration notation:
  - 1<sup>st</sup> – Find the \_\_\_\_\_ that is in the row above the element you want
  - 2<sup>nd</sup> – Write that noble gas's \_\_\_\_\_ in [brackets]
  - 3<sup>rd</sup> – Then continue with the e<sup>-</sup> configuration starting with the next element

- Ex- Scandium: **Long:**  $1s^2 2s^2 2p^6 3s^2 3p^6 4s^2 3d^1$

**Shorthand:**

- Ex- Chlorine: **Long:**  $1s^2 2s^2 2p^6 3s^2 3p^5$

**Shorthand:**

## ELECTRON EXCEPTIONS

- Write the configuration for the following:
  - **Cr:**
  - **Cu:**
- *What they actually are:*
  - **Cr:**
  - **Cu:**
- **Reason** – \_\_\_\_\_ sublevels are the most \_\_\_\_\_.  
\_\_\_\_\_ sublevels are not as stable as filled, but more stable than others.

## ORBITAL NOTATIONS & THE RULES

- **Orbital Notations**
  - Use a \_\_\_\_\_ to represent each \_\_\_\_\_.
    - s orbitals have \_\_\_\_\_ line
    - p orbitals have \_\_\_\_\_ lines
    - d orbitals have \_\_\_\_\_ lines
    - f orbitals have \_\_\_\_\_ lines
  - Use up/down arrows to represent \_\_\_\_\_.
  - Each line can hold a maximum of \_\_\_\_\_ electrons.
  - **Example**  
Titanium: \_\_\_\_\_  
 $1s \quad 2s \quad 2p \quad 3s \quad 3p \quad 4s \quad 3d$

## FILLING RULES

- **Aufbau Principle**-

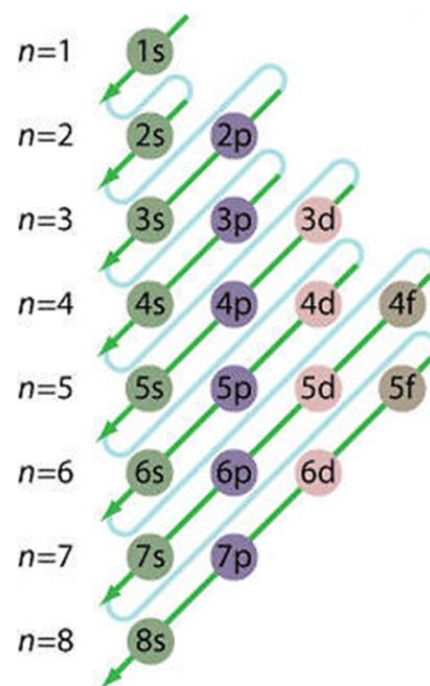
- This is the order we get from “reading” the Periodic Table.

- **Pauli Exclusion Principle**-

- This is the electron “spin.” Either  $+\frac{1}{2}$  or  $-\frac{1}{2}$

- **Hund’s Rule** –

- They’d rather spread out.



## ELECTRON CONFIGURATION PRACTICE

- Sulfur (S)
- Cobalt (Co)
- Strontium (Sr)
- Molybdenum (Mo)
- Antimony (Sb)
- Chlorine (Cl)
- Calcium (Ca)
- Chromium (Cr)
- Zinc (Zn)
- Selenium (Se)
- Mercury (Hg)
- $1s^2 2s^2 2p^6 3s^2 3p^4$
- $1s^2 2s^2 2p^6 3s^2 3p^6 4s^2 3d^{10}$
- $1s^2 2s^2 2p^6 3s^2 3p^6 4s^2 3d^{10} 4p^1$

## ELECTRON IN ATOMS

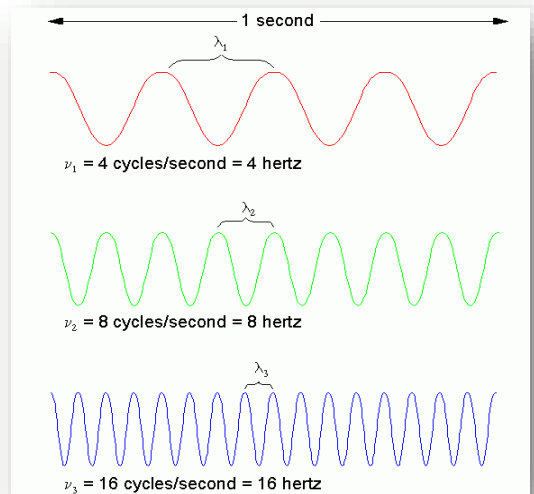
- **Electron Demonstration** – *try to identify key concepts for electrons moving within an atom...*
- **Flame Tests**
  - Elements give off characteristic \_\_\_\_\_ which can be used to identify them.
  - Electrons \_\_\_\_\_ energy from the flame (or other heat source).
  - When a certain amount of energy is reached, which is called a \_\_\_\_\_, electrons jump to a higher energy level called the \_\_\_\_\_.
  - When the electrons \_\_\_\_\_ energy in the form of \_\_\_\_\_, this is also called a **photon** or unit of light, they fall back to the lowest, most \_\_\_\_\_ energy level called the \_\_\_\_\_.

## VISIBLE LIGHT & THE EM SPECTRUM

- Visible light exists as a narrow band of \_\_\_\_\_ that our eyes can detect.
  - The colors of the rainbow \_\_\_\_\_.
  - **Red light** has a wavelength of about \_\_\_\_\_ nanometers and represents \_\_\_\_\_ frequencies.
  - **Violet light** has a wavelength of about \_\_\_\_\_ nanometers and represents \_\_\_\_\_ frequencies.

- **Wave Statistics**

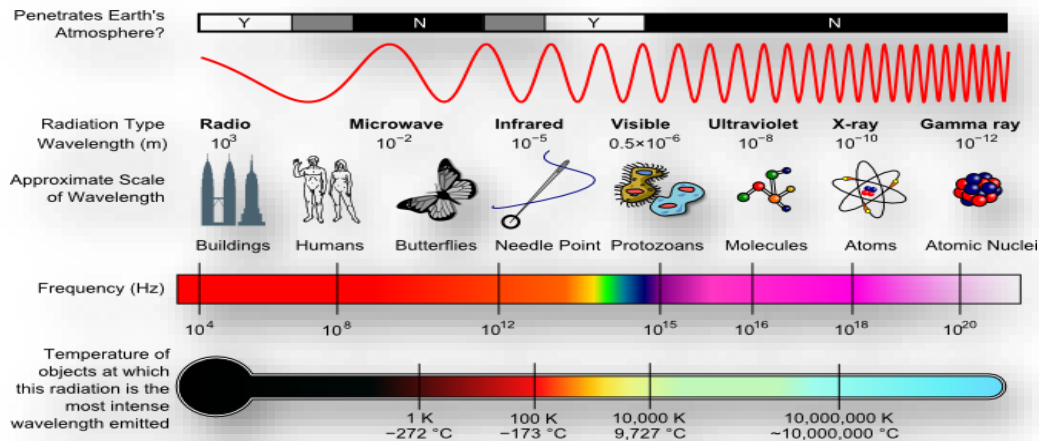
- **Amplitude** – the \_\_\_\_\_ of the wave from zero to crest.
- **Wavelength** – the distance between \_\_\_\_\_ points in phase.
  - Unit:
  - Symbol:
- **Frequency** – the number of cycles (wave peaks) that occur in a unit of time.
  - Unit:
  - Symbol:
- Wavelength & frequency are \_\_\_\_\_ related, meaning that \_\_\_\_\_ wavelengths go with \_\_\_\_\_ frequencies and \_\_\_\_\_ wavelengths go with \_\_\_\_\_ frequencies.



- **Wave Equation**

- Formula:
- Speed of light is always:
- **Example 1** - If the frequency of radiation for yellow light is  $5.10 \times 10^{14}$  Hz, what is the wavelength?
- **Example 2** - What is the frequency of radiation with a wavelength of  $5.00 \times 10^{-8}$  m?

- **Electromagnetic Spectrum** – a grouping of all waves that travel at the speed of light.

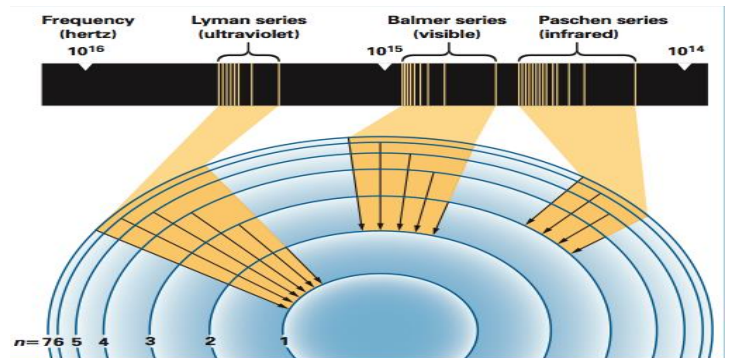


### ATOMIC EMISSION SPECTRA

- Electrons returning from an \_\_\_\_\_ energy level emit \_\_\_\_\_ of specific \_\_\_\_\_ (specific bands of color).
- Each element has a \_\_\_\_\_ emission spectra and therefore is a good way to \_\_\_\_\_ an element!



- Additional transitions exist, but we can't see them because our eyes only detect visible light. These series of transitions are called Lyman, Balmer and Paschen.



- **Energy** – as an electron falls from excited states they release a **quantum** of energy that can be calculated using Plank's constant and the frequency of the transition.
  - *We will not be calculating these energy transitions in this class.*